

SINGLE SLIT DIFFRACTION

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Reference: Carroll, Bradley W. & Ostlie, Dale A. (2007), *An Introduction to Modern Astrophysics*, 2nd Edition; Pearson Education - Chapter 6, Problem 6.6.

When monochromatic light passes through a single slit of finite width and is projected on a screen some distance from the slit, a diffraction pattern appears. The analysis of this pattern is similar to that for a diffraction grating, except this time we're dealing with only a single slit, and the slit has a non-zero width.

We'll start by dividing the slit in half. Now suppose we choose a location on the screen such that a ray from the top of the slit travels exactly half a wavelength $\lambda/2$ further to reach the screen than a ray from the middle of the slit. In this case, the two rays interfere destructively, cancelling each other out. Now consider a ray from a distance ϵ below the top of the slit. It will travel a distance $\lambda/2$ farther to reach the screen than a ray starting from $\frac{D}{2} + \epsilon$. We can continue the argument for all values of ϵ from 0 up to $\frac{D}{2}$, so the point on the screen that we've been considering will get no illumination, as all the rays from the slit cancel in pairs.

Suppose we divide up the width D of the slit into an even number $2m$ of segments. Then we can repeat the above argument for a different point on the screen such that a ray from the first segment travels $\lambda/2$ further to reach the screen than a ray in the second segment, and so on, pairing the third and fourth segments, fifth and sixth, etc. Thus for each $m = 1, 2, 3, \dots$ we get a dark fringe on the screen. [Note that the argument doesn't work for an odd number of segments, since the rays from the last segment can't be paired with another segment to produce destructive interference.]

We can find the points on the screen in terms of the angle θ between the direction of the rays hitting the screen and the normal to the screen. [Although rays from different locations in the slit travel at slightly different angles to reach the same point on the screen, in practice, the slit width D is much less than the distance to the screen, so the rays are almost parallel.] By drawing a diagram (see Fig. 6.7 in Carroll & Ostlie) we see that the path difference is $\frac{D}{2m} \sin \theta$ so the condition for a dark fringe is

$$(0.1) \quad \frac{D}{2m} \sin \theta = \frac{\lambda}{2}$$

$$(0.2) \quad \sin \theta = m \frac{\lambda}{D}$$

By a process of adding up wave amplitudes from across the slit, we can work out a formula for the intensity of light at angle θ :

$$(0.3) \quad I(\theta) = I_0 \frac{\sin^2 \gamma}{\gamma^2}$$

where

$$(0.4) \quad \gamma \equiv \frac{\pi D}{\lambda} \sin \theta$$

[This formula bears a superficial resemblance to the intensity of light from a diffraction grating with N slits: $I = I_0 \frac{\sin^2 N\gamma}{\sin^2 \gamma}$. However, the diffraction grating formula ignores the width of each individual slit, so setting $N = 1$ in that formula gives $I = I_0$ with no diffraction pattern at all.]

Because $\lim_{\gamma \rightarrow 0} \frac{\sin \gamma}{\gamma} = 1$, $I_0 = I(0)$. This can also be done using l'Hôpital's rule:

$$(0.5) \quad \lim_{\gamma \rightarrow 0} \frac{\sin \gamma}{\gamma} = \lim_{\gamma \rightarrow 0} \frac{\cos \gamma}{1} = 1$$

To get a feel for the spacing of the fringes, suppose the slit has a width of $D = 1.0 \mu\text{m}$ and $\lambda = 500 \text{ nm}$. Then from 0.2 with $m = 1$ we have

$$(0.6) \quad \sin \theta = \frac{5 \times 10^{-7}}{10^{-6}} = 0.5$$

$$(0.7) \quad \theta = 30^\circ$$

In this case, the central peak is quite broad. In practice, if we were observing two light sources through such a slit, they would need to be separated by more than 30° to be resolvable on the screen. Increasing the slit width improves the resolution as it decreases the angle θ at which the minimum occurs, and thus it reduces the width of the central peak, giving a sharper image.

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