

FINAL FORM FOR THE SCALAR QUANTUM FIELD

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From the assumptions that the scalar quantum field must be a linear combination of creation and annihilation operators, and that it must transform properly under Lorentz transformations and translations, we arrived at a form for the scalar quantum field as follows:

$$\phi(x) = \int \frac{d^3\mathbf{p}}{(2\pi)^3 2\omega_{\mathbf{p}}} \left[f e^{-ip \cdot x} \alpha(p) + g e^{ip \cdot x} \alpha^\dagger(p) \right] \quad (1)$$

We still need to satisfy what might be seen as the two most important requirements. First, the field must be hermitian, since it's an observable, and second, the field operators $\phi(x)$ and $\phi(y)$ must be independent of each other (that is, they must commute) if the spacetime interval $x - y$ is spacelike, so that $(x - y)^2 < 0$. That is

$$[\phi(x), \phi(y)] = 0 \quad \text{if } (x - y)^2 < 0 \quad (2)$$

To satisfy these conditions, Coleman proposes two independent versions for ϕ in his equation 3.34. He reverts to the a instead of α for the operators, using

$$\begin{aligned} \alpha^\dagger(p) &\equiv (2\pi)^{3/2} \sqrt{2\omega_{\mathbf{p}}} a_{\mathbf{p}}^\dagger \\ \alpha(p) &\equiv (2\pi)^{3/2} \sqrt{2\omega_{\mathbf{p}}} a_{\mathbf{p}} \end{aligned} \quad (3)$$

These use the partial fields

$$\phi^{(+)}(x) = \int \frac{d^3\mathbf{p}}{\sqrt{(2\pi)^3} \sqrt{2\omega_{\mathbf{p}}}} a_{\mathbf{p}} e^{-ip \cdot x} \quad (4)$$

$$\phi^{(-)}(x) = \int \frac{d^3\mathbf{p}}{\sqrt{(2\pi)^3} \sqrt{2\omega_{\mathbf{p}}}} a_{\mathbf{p}}^\dagger e^{ip \cdot x} \quad (5)$$

Note that

$$\phi^{(+)}(x)^\dagger = \phi^{(-)}(x) \quad (6)$$

We then have

$$\phi(x) = f\phi^{(+)}(x) + g\phi^{(-)}(x) \quad (7)$$

In order to satisfy hermiticity, we need

$$\phi^\dagger(x) = \phi(x) \quad (8)$$

from which we see that

$$f = g^\dagger \quad (9)$$

Since any multiple of $\phi(x)$ is a valid field, we might as well take

$$|f| = |g| = 1 \quad (10)$$

This means that the most general form for the field is

$$\phi(x) = e^{i\theta}\phi^{(+)}(x) + e^{-i\theta}\phi^{(-)}(x) \quad (11)$$

where θ is any real number.

Two independent forms for the field are ϕ^1 with $\theta = 0$ and ϕ^2 with $\theta = \frac{\pi}{2}$:

$$\phi^1(x) = \phi^{(+)}(x) + \phi^{(-)}(x) \quad (12)$$

$$\phi^2(x) = i\left(\phi^{(+)}(x) - \phi^{(-)}(x)\right) \quad (13)$$

We now need to evaluate the commutators of these two possibilities to see if 2 is satisfied. We can do this using the commutators of the creation and annihilation operators:

$$\begin{aligned} [a_{\mathbf{p}}, a_{\mathbf{p}'}^\dagger] &= \delta(\mathbf{p} - \mathbf{p}') \\ [a_{\mathbf{p}}, a_{\mathbf{p}'}] &= [a_{\mathbf{p}}^\dagger, a_{\mathbf{p}'}^\dagger] = 0 \end{aligned} \quad (14)$$

The first possibility is that both ϕ^1 and ϕ^2 (and any linear combination of them) represents an observable. In order for that to be possible, all spacelike commutators between linear combinations of ϕ^1 and ϕ^2 must be zero. In particular, we must have

$$[\phi^1(x), \phi^2(y)] = 0 \quad \text{if } (x - y)^2 < 0 \quad (15)$$

Coleman works out this commutator in detail in equations 3.37 and 3.38, and shows that, the commutator comes out to

$$[\phi^1(x), \phi^2(y)] = -i\Delta_+(x - y) - i\Delta_+(y - x) \quad (16)$$

where

$$\Delta_+(x-y) = \int \frac{d^3\mathbf{p}}{(2\pi)^3 (2\omega_{\mathbf{p}})} e^{-ip \cdot (x-y)} \quad (17)$$

The time derivative of Δ_+ comes out to

$$\frac{\partial}{\partial x^0} \Delta_+(x) = -\frac{i}{2} \int \frac{d^3\mathbf{p}}{(2\pi)^3} e^{-ip \cdot x} \quad (18)$$

which is the same integral we looked at earlier using contour integration, and determined that it isn't zero. If a function has a non-zero derivative, then the function itself cannot be identically zero, so using the two independent forms 12 doesn't give us a field that satisfies causality.

The only correct solution is to use ϕ^1 on its own, and we find (Coleman equation 3.42) that (calling ϕ^1 just ϕ):

$$[\phi(x), \phi(y)] = \Delta_+(x-y) - \Delta_+(y-x) \equiv i\Delta(x-y) \quad (19)$$

This *does* turn out to be zero for spacelike intervals, according to the following argument. First, we observe that Δ_+ , and therefore Δ , is Lorentz invariant, since from its definition 17, it's the integral of the Lorentz invariant measure $d^3\mathbf{p}/2\omega_{\mathbf{p}}$ multiplied by the exponential of a Lorentz scalar $p \cdot (x-y)$. A Lorentz invariant quantity can be worked out in any inertial frame. A spacelike interval is one where the two events must always be separated by a non-zero spatial distance, no matter what inertial frame you use. However, the two events can occur in either order (or simultaneously) with regard to time. As shown in footnote 8 in the book, any spacelike interval can be turned into its negative by a proper Lorentz transformation, so $x-y$ and $y-x$ are related by a single proper Lorentz transformation. Thus $\Delta_+(x-y) = \Delta_+(y-x)$ and thus $\Delta(x-y) = 0$.

To see that this is possible, it might help to look at two events, where $e_1 = (t, x, y, z) = (0, 0, 0, 0)$ and $e_2 = (0, 1, 0, 0)$. In the frame in which these events have these coordinates, they are simultaneous but have a non-zero separation in space, so $e_2 - e_1$ is a spacelike interval. From the argument above, it should be possible to find a Lorentz transformation that turns $e_2 - e_1$ into $e_1 - e_2$. The usual Lorentz transformation for a boost is given as

$$\begin{aligned} x' &= \gamma(x - vt) \\ t' &= \gamma(t - vx) \end{aligned} \quad (20)$$

Since $t = 0$ for both events and $\gamma > 0$ for any velocity v , it won't matter what the relative velocity of two frames is, we'll always have $e'_1 = (0, 0, 0, 0)$ and $e'_2 = (-\gamma v, \gamma, 0, 0)$ so that $e'_2 - e'_1$ will never be equal to $e_1 - e_2$. Thus we can't invert a spacelike vector using a boost.

However, remember that an ordinary rotation in space is also a proper Lorentz transformation, and we can turn a vector pointing in the $+x$ direction into an equal and opposite vector in the $-x$ direction by rotating it by π about the z axis (or the y axis).

The same argument doesn't work for two spacetime points with a time-like separation, since there is always a non-zero time (of the same sign) between these two events. In other words, one event can be connected to the other by a light signal, so we can't reverse the temporal order of the events, meaning that we can't transform $x - y$ to $y - x$ by a single Lorentz transformation.

We therefore end up with the final form for our scalar quantum field:

$$\boxed{\phi(x) = \int \frac{d^3\mathbf{p}}{\sqrt{(2\pi)^3} \sqrt{2\omega_{\mathbf{p}}}} \left[a_{\mathbf{p}} e^{-ip \cdot x} + a_{\mathbf{p}}^\dagger e^{ip \cdot x} \right]} \quad (21)$$

This field is often written as the sum of two parts:

$$\phi(x) = \phi^{(+)}(x) + \phi^{(-)}(x) \quad (22)$$

where

$$\phi^{(+)}(x) = \int \frac{d^3\mathbf{p}}{\sqrt{(2\pi)^3} \sqrt{2\omega_{\mathbf{p}}}} a_{\mathbf{p}} e^{-ip \cdot x} \quad (23)$$

$$\phi^{(-)}(x) = \int \frac{d^3\mathbf{p}}{\sqrt{(2\pi)^3} \sqrt{2\omega_{\mathbf{p}}}} a_{\mathbf{p}}^\dagger e^{ip \cdot x} \quad (24)$$

It seems counterintuitive to name the $+$ field as the one with the annihilation operators, but the rationale appears to be that if we apply the Schrödinger energy operator to this field we get

$$i \frac{\partial}{\partial t} \phi^{(+)}(x) = i \frac{\partial}{\partial t} \int \frac{d^3\mathbf{p}}{\sqrt{(2\pi)^3} \sqrt{2\omega_{\mathbf{p}}}} a_{\mathbf{p}} e^{-ip \cdot x} \quad (25)$$

$$= \int \frac{i d^3\mathbf{p}}{\sqrt{(2\pi)^3} \sqrt{2\omega_{\mathbf{p}}}} a_{\mathbf{p}} \frac{\partial}{\partial t} e^{-i\omega_{\mathbf{p}}t + i\mathbf{p} \cdot \mathbf{x}} \quad (26)$$

$$= \int \frac{(+\omega_{\mathbf{p}}) d^3\mathbf{p}}{\sqrt{(2\pi)^3} \sqrt{2\omega_{\mathbf{p}}}} a_{\mathbf{p}} e^{-i\omega_{\mathbf{p}}t + i\mathbf{p} \cdot \mathbf{x}} \quad (27)$$

That is, we get a $+\omega_{\mathbf{p}}$ after applying the operator $i\frac{\partial}{\partial t}$, so we get a positive frequency, hence the + sign. Applying the Schrödinger energy operator to $\phi^{(-)}$ brings in a factor of $-\omega_{\mathbf{p}}$, that is, a negative frequency.

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