

CIRCULAR ORBIT AROUND A SUPERMASSIVE BLACK HOLE

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Reference: Moore, Thomas A., *A General Relativity Workbook*, University Science Books (2013) - Chapter 10; Problem 10.9.

Suppose we are unfortunate enough to be orbiting a supermassive black hole with a mass of 10^6 solar masses. For the sun, $GM = 1.477$ km, so for the black hole, $GM = 1.477 \times 10^6$ km. The radius of the orbit is $r = 10GM = 1.477 \times 10^7$ km. For comparison, the average earth-sun distance is 1.5×10^8 km, so this orbit is about 10 times closer to the black hole than Earth is to the sun. A rather frightening prospect.

From its radius, we can work out the energy and angular momentum per unit mass. From the formula for angular momentum as measured by the orbiting object:

$$\begin{aligned} (1) \quad l^2 &= \frac{r^2 GM}{r - 3GM} \\ (2) \quad &= \frac{100 (GM)^2}{7} \\ (3) \quad l &= \frac{10}{\sqrt{7}} GM \\ (4) \quad &= 5.58 \times 10^6 \text{ km} \end{aligned}$$

The energy is

$$\begin{aligned} (5) \quad \tilde{E} &= -\frac{GM}{2r} \left(1 - \frac{3GM}{r}\right)^{-1} \left(1 - \frac{4GM}{r}\right) \\ (6) \quad &= -\frac{0.6}{20 \times 0.7} \\ (7) \quad &= -0.043 \end{aligned}$$

The period of the orbit as measured by the spacecraft can be found from

$$\begin{aligned} (8) \quad & l = r^2 \omega \\ (9) \quad & = \frac{2\pi r^2}{T} \\ (10) \quad & T = \frac{2\pi r^2}{l} \\ (11) \quad & = \frac{200\sqrt{7}\pi}{10} GM \\ (12) \quad & = 166GM \\ (13) \quad & = 2.455 \times 10^8 \text{ km} \\ (14) \quad & = 821 \text{ s} \\ (15) \quad & = 13.7 \text{ minutes} \end{aligned}$$