

## CIRCULAR ORBIT AROUND A SUPERMASSIVE BLACK HOLE

Link to: [physicspages home page](#).

To leave a comment or report an error, please use the [auxiliary blog](#).

Reference: Moore, Thomas A., *A General Relativity Workbook*, University Science Books (2013) - Chapter 10; Problem 10.9.

Suppose we are unfortunate enough to be orbiting a supermassive black hole with a mass of  $10^6$  solar masses. For the sun,  $GM = 1.477$  km, so for the black hole,  $GM = 1.477 \times 10^6$  km. The radius of the orbit is  $r = 10GM = 1.477 \times 10^7$  km. For comparison, the average earth-sun distance is  $1.5 \times 10^8$  km, so this orbit is about 10 times closer to the black hole than Earth is to the sun. A rather frightening prospect.

From its radius, we can work out the energy and angular momentum per unit mass. From the formula for angular momentum as measured by the orbiting object:

$$\begin{aligned}(0.1) \quad l^2 &= \frac{r^2 GM}{r - 3GM} \\(0.2) \quad &= \frac{100(GM)^2}{7} \\(0.3) \quad l &= \frac{10}{\sqrt{7}} GM \\(0.4) \quad &= 5.58 \times 10^6 \text{ km}\end{aligned}$$

The energy is

$$\begin{aligned}(0.5) \quad \tilde{E} &= -\frac{GM}{2r} \left(1 - \frac{3GM}{r}\right)^{-1} \left(1 - \frac{4GM}{r}\right) \\(0.6) \quad &= -\frac{0.6}{20 \times 0.7} \\(0.7) \quad &= -0.043\end{aligned}$$

The period of the orbit as measured by the spacecraft can be found from

$$\begin{aligned} (0.8) \quad l &= r^2 \omega \\ (0.9) \quad &= \frac{2\pi r^2}{T} \\ (0.10) \quad T &= \frac{2\pi r^2}{l} \\ (0.11) \quad &= \frac{200\sqrt{7}\pi}{10} GM \\ (0.12) \quad &= 166GM \\ (0.13) \quad &= 2.455 \times 10^8 \text{ km} \\ (0.14) \quad &= 821 \text{ s} \\ (0.15) \quad &= 13.7 \text{ minutes} \end{aligned}$$