

EINSTEIN EQUATION FOR A PERFECT FLUID

Link to: physicspages home page.

To leave a comment or report an error, please use the auxiliary blog.

Reference: Moore, Thomas A., *A General Relativity Workbook*, University Science Books (2013) - Chapter 21; Problem 21.3.

The Einstein equation can be written as

$$(0.1) \quad R^{ij} = \kappa \left(T^{ij} - \frac{1}{2} g^{ij} T \right) + \Lambda g^{ij}$$

For a perfect fluid the stress-energy tensor is

$$(0.2) \quad T^{ij} = (\rho_0 + P_0) u^i u^j + P_0 g^{ij}$$

where u^i is the four-velocity, ρ_0 is the energy density and P_0 is the pressure. We have for the stress-energy scalar:

$$(0.3) \quad T = (\rho_0 + P_0) g_{ij} u^i u^j + P_0 g_{ij} g^{ij}$$

$$(0.4) \quad = -(\rho_0 + P_0) + 4P_0$$

$$(0.5) \quad = 3P_0 - \rho_0$$

since $g_{ij} g^{ij} = 4$ and $g_{ij} u^i u^j = -1$ in any coordinate system.

The stress energy term in 0.1 is therefore

$$(0.6) \quad T^{ij} - \frac{1}{2} g^{ij} T = (\rho_0 + P_0) u^i u^j + P_0 g^{ij} - \frac{1}{2} g^{ij} (3P_0 - \rho_0)$$

$$(0.7) \quad = (\rho_0 + P_0) u^i u^j + \frac{1}{2} g^{ij} (\rho_0 - P_0)$$

In a local orthogonal frame (LOF) in which the fluid is at rest $u^t = 1$, $u^i = 0$ for $i = x, y, z$ and $g^{ij} = \eta^{ij}$, so

$$(0.8) \quad T^{tt} - \frac{1}{2}g^{tt}T = \rho_0 + P_0 - \frac{1}{2}(\rho_0 - P_0)$$

$$(0.9) \quad = \frac{1}{2}(\rho_0 + 3P_0)$$

$$(0.10) \quad T^{xx} - \frac{1}{2}g^{xx}T = \frac{1}{2}(\rho_0 - P_0)$$

$$(0.11) \quad T^{yy} - \frac{1}{2}g^{yy}T = \frac{1}{2}(\rho_0 - P_0)$$

$$(0.12) \quad T^{zz} - \frac{1}{2}g^{zz}T = \frac{1}{2}(\rho_0 - P_0)$$

PINGBACKS

Pingback: Einstein equation solution for the interior of a spherically symmetric star

Pingback: Cosmic strings