

## ENSEMBLE OF ELECTRONS IN MAGNETIC FIELD

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Shankar, R. (1994), *Principles of Quantum Mechanics*, Plenum Press.

Exercise 14.4.6.

[If some equations are too small to read easily, use your browser's magnifying option (Ctrl + on Chrome, probably something similar on other browsers).]

As an example of a density matrix, we can apply it to an ensemble of spin  $\frac{1}{2}$  particles. The density matrix is defined as

$$(1) \quad \rho \equiv \sum_i p_i |i\rangle \langle i|$$

where  $p_i$  is the probability of a single system being state  $|i\rangle$ . For a spin  $\frac{1}{2}$  particle, there are only 2 states, so the density matrix can be written as a  $2 \times 2$  matrix, once we define a basis for the states (for example, the basis of  $S_z$  states where  $S_z = \pm \frac{\hbar}{2}$ ). Since any  $2 \times 2$  matrix can be written as a linear combination of the Pauli matrices and the identity matrix, the density matrix can be written as

$$(2) \quad \rho = a_0 I + \mathbf{A} \cdot \boldsymbol{\sigma}$$

Since the trace of each of the Pauli matrices is zero, and  $\text{Tr} I = 2$ , we have

$$(3) \quad \text{Tr} \rho = 2a_0$$

However, we know that  $\text{Tr} \rho = 1$ , so we must have  $a_0 = \frac{1}{2}$ , so we can write

$$(4) \quad \rho = \frac{1}{2} (I + \mathbf{a} \cdot \boldsymbol{\sigma})$$

for some vector  $\mathbf{a}$  whose elements are complex numbers.

To find the average value  $\langle \bar{\Omega} \rangle$  of an observable  $\Omega$  in an ensemble, we can use the density matrix in the form

$$(5) \quad \langle \bar{\Omega} \rangle = \text{Tr} (\Omega \rho)$$

To find  $\langle \vec{\sigma} \rangle$ , we can work out each component separately. For  $\sigma_x$  we have, using the properties of the  $\sigma_i$ :

$$\begin{aligned}
 (6) \quad \langle \vec{\sigma}_x \rangle &= \text{Tr}(\sigma_x \rho) \\
 (7) \quad &= \frac{1}{2} \text{Tr}(\sigma_x I + a_x \sigma_x^2 + a_y \sigma_x \sigma_y + a_z \sigma_x \sigma_z) \\
 (8) \quad &= \frac{1}{2} \text{Tr}(\sigma_x + a_x I + i a_y \sigma_z - i a_z \sigma_y) \\
 (9) \quad &= \frac{1}{2} (0 + 2a_x + 0 + 0) \\
 (10) \quad &= a_x
 \end{aligned}$$

We can do similar calculations to get the other two components, with the result

$$(11) \quad \langle \vec{\sigma} \rangle = \mathbf{a}$$

Finally, suppose we have an ensemble of electrons in a constant magnetic field  $\mathbf{B} = B\hat{\mathbf{z}}$ , and that this ensemble is in thermal equilibrium at temperature  $T$ . A central result of statistical mechanics (which we haven't covered yet) is that particles in thermal equilibrium obey the Boltzmann distribution, which states that the probability of finding a particle with energy  $E$  in the ensemble is

$$(12) \quad p_E \propto e^{-E/kT}$$

where  $k$  is the Boltzmann constant. In this case, the energy is that of a magnetic moment  $\mu$  in a constant magnetic field  $\mathbf{B}$ , which is

$$(13) \quad H = -\mu \cdot \mathbf{B} = -\gamma \mathbf{S} \cdot \mathbf{B} = -\gamma S_z B$$

There are only two states ( $S_z = \pm \frac{\hbar}{2}$ ), so the probabilities are

$$(14) \quad p_{\uparrow} = \frac{1}{P} e^{\gamma B \hbar / 2kT}$$

$$(15) \quad p_{\downarrow} = \frac{1}{P} e^{-\gamma B \hbar / 2kT}$$

$$(16) \quad P = e^{\gamma B \hbar / 2kT} + e^{-\gamma B \hbar / 2kT}$$

The density matrix is therefore

$$(17) \quad \rho = \frac{1}{P} \left( e^{\gamma B \hbar / 2kT} |\uparrow\rangle \langle \uparrow| + e^{-\gamma B \hbar / 2kT} |\downarrow\rangle \langle \downarrow| \right)$$

In the  $S_z$  basis, this is

$$(18) \quad \rho = \frac{1}{P} \begin{bmatrix} e^{\gamma B \hbar / 2kT} & 0 \\ 0 & e^{-\gamma B \hbar / 2kT} \end{bmatrix}$$

We can work out the average magnetic moment for the ensemble as

$$(19) \quad \langle \bar{\mu} \rangle = \text{Tr}(\mu \rho)$$

$$(20) \quad = \frac{\hat{\mathbf{z}}}{P} \left[ \frac{\gamma \hbar}{2} e^{\gamma B \hbar / 2kT} - \frac{\gamma \hbar}{2} e^{-\gamma B \hbar / 2kT} \right]$$

$$(21) \quad = \frac{e^{\gamma B \hbar / 2kT} - e^{-\gamma B \hbar / 2kT}}{e^{\gamma B \hbar / 2kT} + e^{-\gamma B \hbar / 2kT}} \frac{\gamma \hbar}{2} \hat{\mathbf{z}}$$

$$(22) \quad = \frac{\gamma \hbar}{2} \tanh \frac{\gamma B \hbar}{2kT} \hat{\mathbf{z}}$$

$$(23) \quad = -\frac{e \hbar}{2mc} \tanh \left( -\frac{e B \hbar}{2mckT} \right) \hat{\mathbf{z}}$$

$$(24) \quad = \frac{e \hbar}{2mc} \tanh \frac{e B \hbar}{2mckT} \hat{\mathbf{z}}$$