

LOCALLY FLAT COORDINATES ON THE SPHERE

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References: Anthony Zee, *Einstein Gravity in a Nutshell*, (Princeton University Press, 2013) - Chapter I.6, Exercise 3.

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Using a pair of coordinates denoted (κ, ζ) , we are given a metric which is, up to first order in the coordinates

$$g_{\kappa\kappa} = 1 \quad (1)$$

$$g_{\zeta\zeta} = 1 + 2\kappa \quad (2)$$

$$g_{\kappa\zeta} = 0 \quad (3)$$

If we make the transformation

$$\kappa = \omega + \frac{1}{2}\phi^2 \quad (4)$$

$$\zeta = \phi - \omega\phi \quad (5)$$

we can get a metric with no linear terms. We need to use the metric transformation formula given as Zee's equation (8) in chapter I.6:

$$g'_{\lambda\sigma}(x') = g_{\mu\nu}(x) \frac{\partial x^\mu}{\partial x'^\lambda} \frac{\partial x^\nu}{\partial x'^\sigma} \quad (6)$$

We therefore need the derivatives:

$$\frac{\partial \kappa}{\partial \omega} = 1 \quad (7)$$

$$\frac{\partial \kappa}{\partial \phi} = \phi \quad (8)$$

$$\frac{\partial \zeta}{\partial \omega} = -\phi \quad (9)$$

$$\frac{\partial \zeta}{\partial \phi} = 1 - \omega \quad (10)$$

We then have

$$g'_{\omega\omega} = 1 + (1 + 2\kappa) \phi^2 \quad (11)$$

$$= 1 + \phi^4 + (2\omega + 1) \phi^2 \quad (12)$$

$$g'_{\omega\phi} = \phi + (1 + 2\kappa) \phi (\omega - 1) \quad (13)$$

$$= 2\phi\omega^2 + (\phi^3 - \phi) \omega - \phi^3 \quad (14)$$

$$g'_{\phi\phi} = \phi^2 + (1 + 2\kappa) (1 - \omega)^2 \quad (15)$$

$$= 1 + 2\omega^3 + (\phi^2 - 3) \omega^2 - 2\omega\phi^2 + 2\phi^2 \quad (16)$$

Although the new metric isn't pretty, we can see that there are no terms linear in either ω or ϕ .

We can now relate this to the flat metric in polar coordinates

$$ds^2 = dr^2 + r^2 d\theta^2 \quad (17)$$

We assume that we start at some point $P = (r, \theta) = (r_*, 0)$. Then if we move a small distance away from this point and take (κ, ζ) to be the deviation we have

$$\kappa = r - r_* \quad (18)$$

$$\zeta = \theta \quad (19)$$

where we're assuming that r is close to r_* and θ is small. Then we have

$$ds^2 = dr^2 + (\kappa + r_*)^2 d\theta^2 \quad (20)$$

$$= d\kappa^2 + (r_*^2 + 2\kappa r_* + \kappa^2) d\zeta^2 \quad (21)$$

$$= d\kappa^2 + \left(1 + \frac{2\kappa}{r_*} + \frac{\kappa^2}{r_*^2}\right) (r_* d\zeta)^2 \quad (22)$$

$$= r_*^2 \left[\frac{d\kappa^2}{r_*^2} + \left(1 + \frac{2\kappa}{r_*} + \frac{\kappa^2}{r_*^2}\right) d\zeta^2 \right] \quad (23)$$

If we scale $\kappa \rightarrow \frac{\kappa}{r_*}$ and $\zeta \rightarrow r_* \zeta$ this is equivalent to

$$ds^2 = d\kappa^2 + (\kappa + 1)^2 d\zeta^2 \quad (24)$$

Thus the metric is the same as the flat metric in polar coordinates with $\kappa + 1$ playing the role of r and ζ playing the role of θ . Since we're taking κ to be very small, we can assume that $\kappa \ll 1$. Also, we're assuming θ is the deviation from 0 angle.